

QM3

Quantum Mechanics 3 notes by [K. Sreeman Reddy](#).

1. [The Klein-Gordon equation](#)
 1. [Electromagnetic interaction](#)
 2. [Conserved current](#)
 3. [Non-relativistic limit](#)
2. [The Dirac equation](#)
3. [Classical Field Theory](#)
 1. [Introduction](#)
 1. [A vibrating system](#)
 1. [Increase the degrees of freedom to N from 2](#)
 2. [Quantizing it](#)
 2. [Classical string](#)
 1. [Quantizing it](#)
 2. [Lagrange formulation](#)
 3. [Hamilton formulation](#)
 4. [Noether's theorem](#)
 1. [Conservation laws](#)
4. [The free scalar quantum field](#)
 1. [Canonical or 2nd Quantization](#)
 2. [Heisenberg picture](#)
 3. [Fourier decomposition of the field](#)
 4. [Creation and annihilation operators](#)
 5. [Wavefunctions](#)
 1. [Normalisation](#)
 6. [Causality](#)
5. [Interacting Scalar Fields](#)

We will use the metric signature $\eta_{\mu\nu} = \text{diag}(+1, -1, -1, -1)$

The Klein-Gordon equation

$$(\square + \left(\frac{mc}{\hbar}\right)^2)\psi(x) = 0, \quad \text{where } \square = \partial^\mu \partial_\mu$$

| | Position space $x = (ct, \mathbf{x})$ | Fourier transformation $\omega = E/\hbar, \quad \mathbf{k} = \mathbf{p}/\hbar$ | Momentum space $p = (E/c, \mathbf{p})$ |
|--------------------------|---|--|---|
| Separated time and space | $\left(\frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \nabla^2 + \frac{m^2 c^2}{\hbar^2}\right) \psi(t, \mathbf{x}) = 0$ | $\psi(t, \mathbf{x}) = \int \frac{d\omega}{2\pi\hbar} \int \frac{d^3k}{(2\pi\hbar)^3} e^{-i(\omega t - \mathbf{k}\cdot\mathbf{x})} \psi(\omega, \mathbf{k})$ | $E^2 = \mathbf{p}^2 c^2 + m^2 c^4$ |
| Four-vector form | $(\square + \left(\frac{mc}{\hbar}\right)^2)\psi(x) = 0,$ | $\psi(x) = \int \frac{d^4p}{(2\pi)^4} e^{-ip\cdot x} \psi(p)$ | $p^2 = +m^2$ |

Electromagnetic interaction

Let $D_\mu = \partial_\mu + \frac{iq}{c} A_\mu$ then under minimal coupling with A_μ

$$D_\mu D^\mu \phi = -(\partial_{ct} + \frac{iq}{\hbar} A_0)^2 \phi + (\partial_i + \frac{iq}{\hbar} A_i)^2 \phi = -\left(\frac{mc}{\hbar}\right)^2 \phi$$

$$(p_\mu - \frac{q}{c} A_\mu)(p^\mu - \frac{q}{c} A^\mu) = m^2 c^2 \phi$$

Conserved current

$$\partial_\mu J^\mu(x) = 0, \quad J^\mu(x) \equiv \frac{i\hbar}{2m} (\phi^*(x)\partial^\mu\phi(x) - \phi(x)\partial^\mu\phi^*(x)) - \frac{q}{mc} A^\mu\phi^*\phi$$

Non-relativistic limit

$$\psi(x, t) = \phi(x, t) e^{-\frac{i}{\hbar}mc^2t} \quad \text{where} \quad \phi(x, t) = u_E(x) e^{-\frac{i}{\hbar}E't}$$

Since $\left| i\hbar \frac{\partial\phi}{\partial t} \right| = E'\phi \ll mc^2\phi$ and $|qA_0\phi| \ll mc^2\phi$ in the non-relativistic limit

$$i\hbar \frac{\partial\phi}{\partial t} = \left[\frac{1}{2m} \left(\vec{p} - \frac{q}{c} \vec{A} \right)^2 + mc^2 + qA^0 \right] \phi$$

we get back the classical Pauli field.

The Dirac equation

$$i\hbar\gamma^\mu\partial_\mu\psi - mc\psi = 0$$

In Dirac representation

$$\gamma^0 = \begin{pmatrix} I_2 & 0 \\ 0 & -I_2 \end{pmatrix}, \quad \gamma^k = \begin{pmatrix} 0 & \sigma^k \\ -\sigma^k & 0 \end{pmatrix}, \quad \gamma^5 = \begin{pmatrix} 0 & I_2 \\ I_2 & 0 \end{pmatrix}$$

In Weyl (chiral) representation

$$\gamma^0 = \begin{pmatrix} 0 & I_2 \\ I_2 & 0 \end{pmatrix}, \quad \gamma^k = \begin{pmatrix} 0 & \sigma^k \\ -\sigma^k & 0 \end{pmatrix}, \quad \gamma^5 = \begin{pmatrix} -I_2 & 0 \\ 0 & I_2 \end{pmatrix}$$

In Weyl (chiral) representation (alternate form)

$$\gamma^0 = \begin{pmatrix} 0 & -I_2 \\ -I_2 & 0 \end{pmatrix}, \quad \gamma^k = \begin{pmatrix} 0 & \sigma^k \\ -\sigma^k & 0 \end{pmatrix}, \quad \gamma^5 = \begin{pmatrix} I_2 & 0 \\ 0 & -I_2 \end{pmatrix}$$

In Majorana representation

$$\gamma^0 = \begin{pmatrix} 0 & \sigma^2 \\ \sigma^2 & 0 \end{pmatrix}, \quad \gamma^1 = \begin{pmatrix} i\sigma^3 & 0 \\ 0 & i\sigma^3 \end{pmatrix}, \quad \gamma^2 = \begin{pmatrix} 0 & -\sigma^2 \\ \sigma^2 & 0 \end{pmatrix},$$

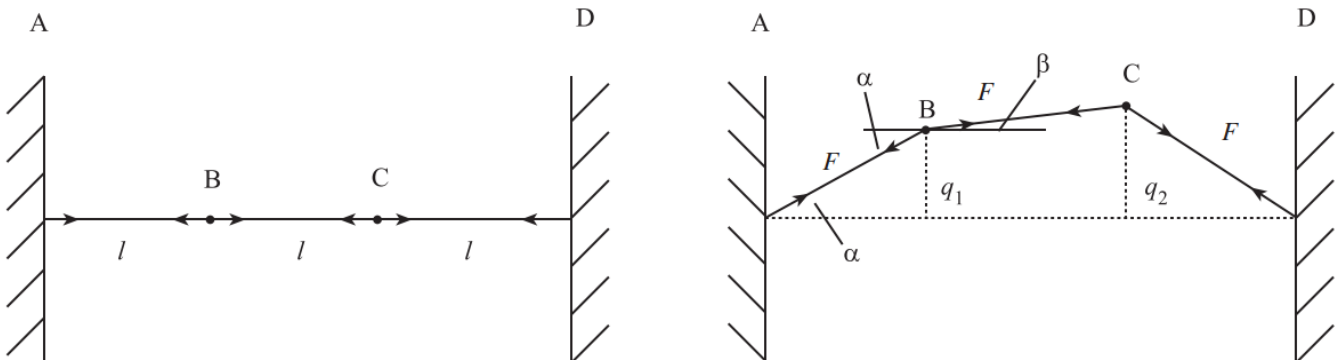
$$\gamma^3 = \begin{pmatrix} -i\sigma^1 & 0 \\ 0 & -i\sigma^1 \end{pmatrix}, \quad \gamma^5 = \begin{pmatrix} \sigma^2 & 0 \\ 0 & -\sigma^2 \end{pmatrix}, \quad C = \begin{pmatrix} 0 & -i\sigma^2 \\ -i\sigma^2 & 0 \end{pmatrix},$$

Classical Field Theory

In a classical string, there are uncountable number of particles but only countable modes **if we fix some boundary conditions**. But if no boundary conditions we will have uncountable modes.

Introduction

A vibrating system



The below 2 equations are linear because we neglected the higher order terms. But quantum mechanics is believed to be **a linear theory without any approximation**.

$$\begin{aligned}
m\ddot{q}_1 &= -\partial V / \partial q_1 \\
m\ddot{q}_2 &= -\partial V / \partial q_2 \\
V &= k (q_1^2 + q_2^2 - q_1 q_2)
\end{aligned}$$

Now observe that if we define

$$Q_1 = (q_1 + q_2) / \sqrt{2} \quad Q_2 = (q_1 - q_2) / \sqrt{2}$$

then for $\omega_1 = \sqrt{\frac{k}{m}}$ and $\omega_2 = \sqrt{\frac{3k}{m}}$

$$\begin{aligned}
m\ddot{Q}_1 &= -\partial V / \partial Q_1 \\
m\ddot{Q}_2 &= -\partial V / \partial Q_2 \\
V \equiv V(Q_1, Q_2) &= \frac{1}{2} m \omega_1^2 Q_1^2 + \frac{1}{2} m \omega_2^2 Q_2^2
\end{aligned}$$

A remarkable thing has happened: the two combinations $q_1 + q_2$ and $q_1 - q_2$ of the original coordinates satisfy uncoupled equations. These are called **normal modes** or **modes**.

- In general the system is in 'a superposition of modes'.
- Modes do not interact.
- The simple change of variables $(q_1, q_2) \rightarrow (Q_1, Q_2)$ does remove the $q_1 q_2$ coupling, this would not be the case if, say, cubic terms in V were to be considered.

Increase the degrees of freedom to N from 2

In general we can find the **mode coordinates** or **normal coordinates**

$$Q_r = \sum_{s=1}^N a_{rs} q_s$$

such that

$$\begin{aligned}
E &= \sum_{r=1}^N \frac{1}{2} m \dot{q}_r^2 + V(q_1, \dots, q_r) \\
E &= \sum_{r=1}^N \left[\frac{1}{2} m \dot{Q}_r^2 + \frac{1}{2} m \omega_r^2 Q_r^2 \right]
\end{aligned}$$

Quantizing it

$$E = \sum_{r=1}^N \left(n_r + \frac{1}{2} \right) \hbar \omega_r$$

We forget about the original N degrees of freedom q_1, q_2, \dots, q_N and the original N 'atoms', which indeed are only remembered in the above equation via the fact that there are N different mode frequencies ω_r . Instead we concentrate on the quanta and treat them as 'things' which really determine the behaviour of our quantum system.

For the state characterized by (n_1, n_2, \dots, n_N) there are n_1 quanta of mode 1 (frequency ω_1), n_2 of mode 2, ... and n_N of mode N .

Classical string



Let $N \rightarrow \infty$ and $Na = l$

$$\frac{1}{c^2} \frac{\partial^2 \phi(x, t)}{\partial t^2} = \frac{\partial^2 \phi(x, t)}{\partial x^2}$$

$$\phi(x, t) = \sum_{r=1}^{\infty} A_r(t) \sin\left(\frac{r\pi x}{\ell}\right)$$

$$E = \int_0^{\ell} \left[\frac{1}{2}\rho \left(\frac{\partial\phi}{\partial t}\right)^2 + \frac{1}{2}\rho c^2 \left(\frac{\partial\phi}{\partial x}\right)^2 \right] dx$$

$$E = (\ell/2) \sum_{r=1}^{\infty} \left[\frac{1}{2}\rho \dot{A}_r^2 + \frac{1}{2}\rho \omega_r^2 A_r^2 \right]$$

Quantizing it

$$E = \sum_{r=1}^{\infty} \left(n_r + \frac{1}{2} \right) \hbar \omega_r$$

We remark that as $\ell \rightarrow \infty$, the mode sum will be replaced by an integral over a continuous frequency variable.

Lagrange formulation

$$\frac{\partial \mathcal{L}}{\partial \phi^a} - \partial_{\mu} \left(\frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi^a)} \right) = 0$$

$$S = \int dt \int d^3x \mathcal{L} = \int d^4x \mathcal{L}$$

$$\mathcal{L}' = \mathcal{L} + \partial_{\mu} K^{\mu}(\Phi)$$

The above Lagrangian also is equivalent to \mathcal{L} .

Hamilton formulation

For many fields $\phi_i(\mathbf{x}, t)$ and their conjugates $\pi_i(\mathbf{x}, t)$ the Hamiltonian density is a function of them all:

$$\mathcal{H}(\phi_1, \phi_2, \dots, \pi_1, \pi_2, \dots, \mathbf{x}, t) = \sum_i \dot{\phi}_i \pi_i - \mathcal{L}(\phi_1, \phi_2, \dots, \nabla \phi_1, \nabla \phi_2, \dots, \partial \phi_1 / \partial t, \partial \phi_2 / \partial t, \dots, \mathbf{x}, t).$$

where each conjugate field is **defined** with respect to its field,

$$\pi_i(\mathbf{x}, t) = \frac{\partial \mathcal{L}}{\partial \dot{\phi}_i}$$

$$H = \int \mathcal{H} d^3x$$

Hamiltonian field equations:

$$\dot{\phi}_i = + \frac{\delta \mathcal{H}}{\delta \pi_i}, \quad \dot{\pi}_i = - \frac{\delta \mathcal{H}}{\delta \phi_i}$$

$$\frac{\delta}{\delta \phi_i} = \frac{\partial}{\partial \phi_i} - \partial_{\mu} \frac{\partial}{\partial (\partial_{\mu} \phi_i)}$$

Noether's theorem

If $\Delta \mathcal{L}$, $\Delta \phi$ and Δx are defined by $\mathcal{L}(x) \rightarrow \mathcal{L}(x) + \epsilon \Delta \mathcal{L}$ (we know that $\Delta \mathcal{L} = \partial_{\mu} K^{\mu}$ for some K^{μ}), $\phi \rightarrow \phi + \epsilon \Delta \phi$ and $x \rightarrow x + \epsilon \Delta x$ then the conserved current defined by

$$j^{\mu} = \frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi)} \Delta \phi - K^{\mu}$$

satisfies $\partial_{\mu} j^{\mu} = 0$. (j^{μ} is unique up to a multiplicative constant)

If the symmetry involves more than one field, the conserved current is

$$j^{\mu} = \sum_i \frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi_i)} \Delta \phi_i - K^{\mu}$$

The above conservation law implies that the **Noether charge** over all space is conserved

$$Q = \int d^3 \mathbf{x} j^0$$

is constant in time: $\frac{dQ}{dt} = 0$.

OR:

$$T^{\mu\nu} = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \Phi^A)} \partial^\nu \Phi^A - \eta^{\mu\nu} \mathcal{L}$$

which is called the stress-energy tensor. We can now define a current

$$J^\mu = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \Phi^A)} \delta \Phi^A - T^{\mu\nu} \delta x_\nu$$

and write

$$\delta \mathcal{A} = \int_\Omega d^4 x \partial_\mu J^\mu$$

This relation holds for arbitrary variations of the fields and coordinates provided the equations of motion are satisfied.

Conservation laws

The 1st 4 symmetries together form the Poincaré group

| Conservation Law | Respective Noether symmetry invariance | Number of dimensions |
|--|--|--|
| Conservation of mass-energy | Time-translation invariance | 1: translation along time axis |
| Conservation of linear momentum | Space-translation invariance | 3: translation along x,y,z directions |
| Conservation of angular momentum | Rotation invariance | 3: rotation about x,y,z axes |
| Conservation of CM (center-of-momentum) velocity | Lorentz-boost invariance | 3: Lorentz-boost along x,y,z directions |
| Conservation of electric charge | $U(1)$ Gauge invariance | $1 \otimes 4$: scalar field (1D) in 4D spacetime (x,y,z + time evolution) |
| Conservation of color charge | $SU(3)$ Gauge invariance | 3: r,g,b |
| Conservation of weak isospin | $SU(2)_L$ Gauge invariance | 1: weak charge |
| Conservation of probability | Probability invariance | $1 \otimes 4$: total probability always = 1 in whole x,y,z space, during time evolution |

The free scalar quantum field

$$\mathcal{L}_{\text{KG}} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2$$

Canonical or 2nd Quantization

$$[\hat{\phi}(\vec{x}, t), \hat{\pi}(\vec{y}, t)] = i \delta^3(\vec{x} - \vec{y})$$

Heisenberg picture

$$\dot{\hat{\phi}}(x) = i[\hat{H}, \hat{\phi}(x)]$$

$$\dot{\hat{\pi}}(x) = i[\hat{H}, \hat{\pi}(x)]$$

Fourier decomposition of the field

The four-dimensional analogue of the Fourier expansion of the field ϕ takes the form

$$\hat{\phi}(x) = \int_{-\infty}^{\infty} \frac{d^3 k}{(2\pi)^3 \sqrt{2\omega}} [\hat{a}(k) e^{-ik \cdot x} + \hat{a}^\dagger(k) e^{ik \cdot x}]$$

with a similar expansion for the **conjugate momentum** $\hat{\pi} = \dot{\hat{\phi}}$:

$$\hat{\pi}(x) = \int_{-\infty}^{\infty} \frac{d^3 k}{(2\pi)^3 \sqrt{2\omega}} (-i\omega) [\hat{a}(k)e^{-ik \cdot x} - \hat{a}^\dagger(k)e^{ik \cdot x}]$$

Here $k \cdot x$ is the four-dimensional dot product $k \cdot x = \omega t - \mathbf{k} \cdot \mathbf{x}$, and $\omega = +(\mathbf{k}^2 + m^2)^{1/2}$. A **positive-frequency** solution of the field equation has as its coefficient the operator that **destroys** a particle in that single-particle wavefunction. A **negative-frequency** solution of the field equation, being the Hermitian conjugate of a positive-frequency solution, has as its coefficient the operator that **creates** a particle in that positive-energy single-particle wavefunction.

$$\hat{n}(k) = \hat{a}^\dagger(k)\hat{a}(k)$$

The Hamiltonian is found to be

$$\hat{H}_{\text{KG}} = \int d^3 x \hat{\mathcal{H}}_{\text{KG}} = \int_{-\infty}^{\infty} d^3 x \frac{1}{2} [\hat{\pi}^2 + \nabla \hat{\phi} \cdot \nabla \hat{\phi} + m^2 \hat{\phi}^2]$$

and this can be expressed in terms of the \hat{a} 's and the \hat{a}^\dagger 's using the expansion for $\hat{\phi}$ and $\hat{\pi}$ and the commutator

$$[\hat{a}(k), \hat{a}^\dagger(k')] = (2\pi)^3 \delta^3(\mathbf{k} - \mathbf{k}')$$

with all others vanishing. The result is, as expected,

$$\hat{H}_{\text{KG}} = \frac{1}{2} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} [\hat{a}^\dagger(k)\hat{a}(k) + \hat{a}(k)\hat{a}^\dagger(k)] \omega$$

and, **normally ordering** (operators rearranged with all creation operators on the left) as usual, we arrive at

$$\begin{aligned} \hat{H}_{\text{KG}} &= \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \hat{a}^\dagger(k)\hat{a}(k)\omega \\ \vec{P} &= - \int d^3 x \pi \vec{\nabla} \phi = \int \frac{d^3 p}{(2\pi)^3} \vec{p} a_p^\dagger a_p \end{aligned}$$

Creation and annihilation operators

$$a_{\mathbf{p}} = \frac{i}{\sqrt{2\omega_{\mathbf{p}}}} \int d^3 \mathbf{x} [\Pi(\mathbf{x}) - i\omega_{\mathbf{p}}\phi(\mathbf{x})] e^{-i\mathbf{p} \cdot \mathbf{x}}$$

$$a_{\mathbf{k}}^\dagger = \frac{-i}{\sqrt{2\omega_{\mathbf{k}}}} \int d^3 \mathbf{y} [\Pi(\mathbf{y}) + i\omega_{\mathbf{k}}\phi(\mathbf{y})] e^{i\mathbf{k} \cdot \mathbf{y}}$$

$$[\hat{H}, a_{\mathbf{p}}^\dagger] = \omega_{\mathbf{p}} a_{\mathbf{p}}^\dagger$$

$$[\hat{H}, a_{\mathbf{p}}] = -\omega_{\mathbf{p}} a_{\mathbf{p}}$$

$$\overbrace{[H, [H, [\dots [H, a_{\mathbf{p}}] \dots]]]}^{n \text{ times}} = (-\omega_{\mathbf{p}})^n a$$

$$\overbrace{[\hat{H}, [\hat{H}, [\dots [\hat{H}, a_{\mathbf{p}}^\dagger] \dots]]]}^{n \text{ times}} = (\omega_{\mathbf{p}})^n a$$

$$\begin{aligned} \hat{H} a_{\mathbf{p}} &= a_{\mathbf{p}} (\hat{H} - E_{\mathbf{p}}) \\ \hat{H}^n a_{\mathbf{p}} &= a_{\mathbf{p}} (\hat{H} - E_{\mathbf{p}})^n \end{aligned}$$

$$\begin{aligned} e^{i\hat{H}t} a_{\mathbf{p}} e^{-i\hat{H}t} &= a_{\mathbf{p}} e^{-iE_{\mathbf{p}}t}, \\ e^{i\hat{H}t} a_{\mathbf{p}}^\dagger e^{-i\hat{H}t} &= a_{\mathbf{p}}^\dagger e^{iE_{\mathbf{p}}t} \end{aligned}$$

$$e^{-i\hat{\mathbf{P}} \cdot \mathbf{x}} a_{\mathbf{p}} e^{i\hat{\mathbf{P}} \cdot \mathbf{x}} = a_{\mathbf{p}} e^{i\mathbf{p} \cdot \mathbf{x}}$$

$$e^{-i\hat{\mathbf{P}} \cdot \mathbf{x}} a_{\mathbf{p}}^\dagger e^{i\hat{\mathbf{P}} \cdot \mathbf{x}} = a_{\mathbf{p}}^\dagger e^{-i\mathbf{p} \cdot \mathbf{x}}$$

$$\begin{aligned} \phi(x) &= e^{i(\hat{H}t - \hat{\mathbf{P}} \cdot \mathbf{x})} \phi(0) e^{-i(\hat{H}t - \hat{\mathbf{P}} \cdot \mathbf{x})} \\ &= e^{iP \cdot x} \phi(0) e^{-iP \cdot x} \end{aligned}$$

Wavefunctions

$$\begin{aligned} \langle 0 | \hat{\phi}(x, t) | k' \rangle &= \langle 0 | \int \frac{dk}{2\pi\sqrt{2\omega}} [\hat{a}(k)e^{ikx-i\omega t} + \hat{a}^\dagger(k)e^{-ikx+i\omega t}] N \hat{a}^\dagger(k') | 0 \rangle \\ \langle 0 | \int \frac{N dk}{2\pi\sqrt{2\omega}} [\hat{a}^\dagger(k') \hat{a}(k) + 2\pi\delta(k-k')] e^{ikx-i\omega t} | 0 \rangle &= N \frac{e^{ik'x-i\omega't}}{\sqrt{2\omega'}} \\ \langle 0 | \hat{\phi}(x, t) | 0 \rangle &= 0 \end{aligned}$$

Thus we discover that the vacuum to one-particle matrix elements of the field operators are just the familiar wavefunctions of single-particle quantum mechanics.

$$\langle \mathbf{q} | \phi(\mathbf{x}) | 0 \rangle = \int \frac{d^3\mathbf{p}}{(2\pi)^3} \frac{1}{2E_{\mathbf{p}}} e^{-i\mathbf{p}\cdot\mathbf{x}} \langle \mathbf{q} | \mathbf{p} \rangle = e^{-i\vec{q}\cdot\vec{x}}$$

In the language of second quantization, $e^{i\mathbf{q}\cdot\mathbf{x}}$ tells us how much amplitude there is in the q th momentum mode if we create a scalar particle at spacetime point x .

Normalisation

- $d^4p \delta(p^2 - m^2) \theta(p^0)$ is Lorentz invariant.
- Using $\delta[f(x)] = \sum_{\{x|f(x)=0\}} \frac{1}{|f'(x)|} \delta(x)$ we get

$$\delta(p^2 - m^2) \theta(p_0) = \frac{1}{2E_p} \delta(p_0 - E_p) \theta(p_0)$$

$$|\mathbf{p}\rangle = \sqrt{2E_p} a_{\mathbf{p}}^\dagger |0\rangle$$

$$\langle \mathbf{p} | \mathbf{k} \rangle = 2E_p (2\pi)^3 \delta^{(3)}(\mathbf{p} - \mathbf{k})$$

$$\mathbf{1}_{1\text{-particle}} = \int \frac{d^3\mathbf{p}}{(2\pi)^3} |\mathbf{p}\rangle \frac{1}{2E_p} \langle \mathbf{p}|$$

Causality

If $\phi(x), \phi(y)$ vanishes, one measurement cannot affect the other. In fact, if the commutator vanishes for $(x - y)^2 < 0$, causality is preserved quite generally, since

commutators involving any function of $\phi(x)$, including $\pi(x) = \frac{\partial\phi}{\partial t}$, would also have to vanish.

$$\begin{aligned} [\phi(x), \phi(y)] &= \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \int \frac{d^3q}{(2\pi)^3} \frac{1}{\sqrt{2E_q}} \\ &\times [(a_{\mathbf{p}} e^{-ip\cdot x} + a_{\mathbf{p}}^\dagger e^{ip\cdot x}), (a_{\mathbf{q}} e^{-iq\cdot y} + a_{\mathbf{q}}^\dagger e^{iq\cdot y})] \\ &= \int \frac{d^3p}{(2\pi)^3} \frac{1}{2E_p} (e^{-ip\cdot(x-y)} - e^{ip\cdot(x-y)}) \\ &= D(x - y) - D(y - x) \end{aligned}$$

- $[\phi(x), \phi(y)]$ is Lorentz invariant and it is a complex number.
- When $(x - y)^2 < 0$ we can always change our frame such that $(x - y) \rightarrow -(x - y)$ and the commutator will be 0. This we can't do if $(x - y)^2 > 0$.

In quantum field theory, causality requires that every particle have a corresponding antiparticle with the same mass and opposite quantum numbers (in this case electric charge). For the real-valued Klein-Gordon field, the particle is its own antiparticle.

Interacting Scalar Fields
